

Tour of First PAS Measurements

PAS/SWA team

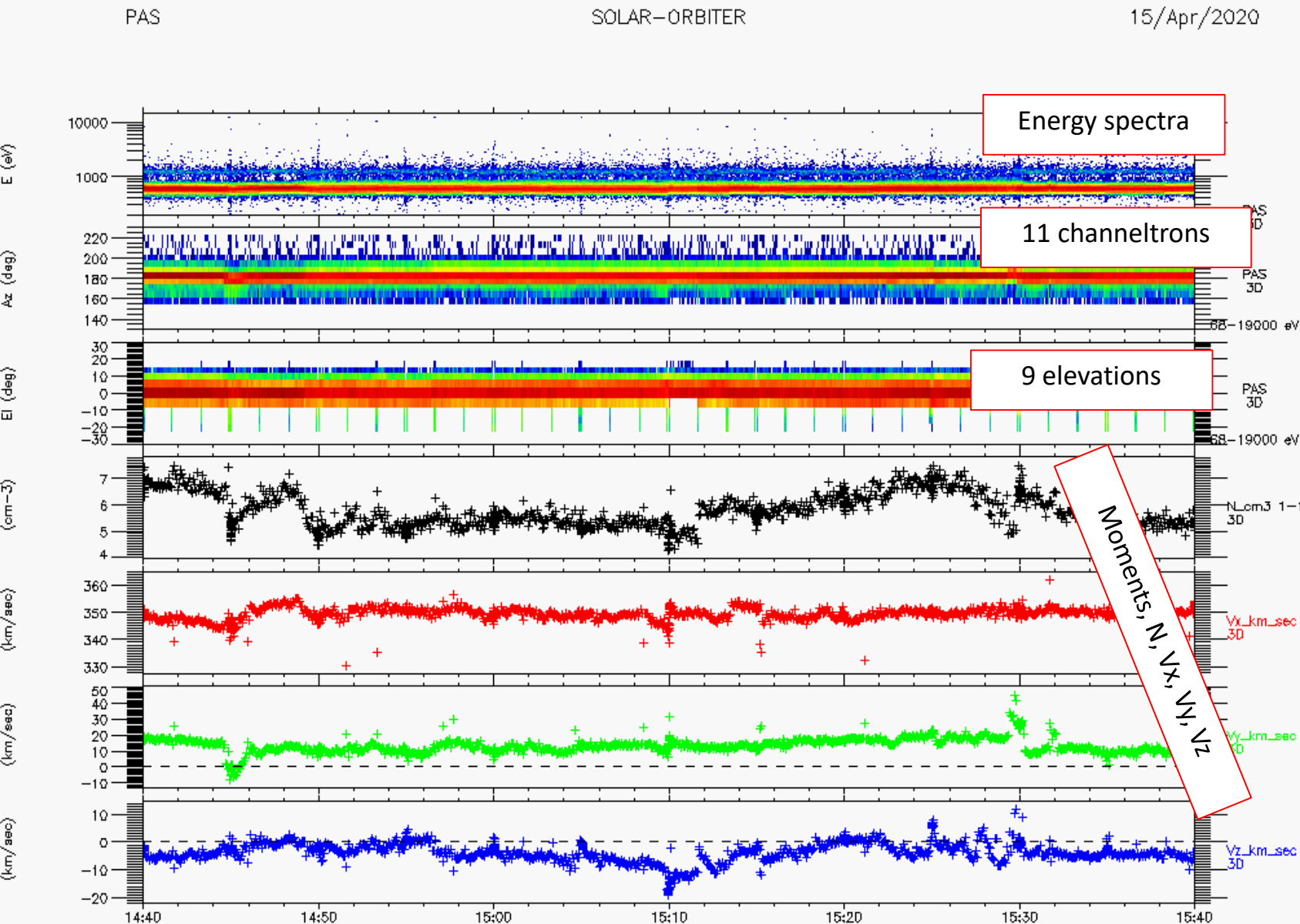


Main points:

- 1) PAS appears to be a very nice tool.*
- 2) We perhaps already see 'things' in SW that was not described before.*
- 3) The L0 -> L2 software is (almost) ready. Scientifically meaningful measurements will be timely delivered to the community.*



PAS commissioning - April, 14 th and 15 th – First data



First hour of PAS working in nominal scientific mode.

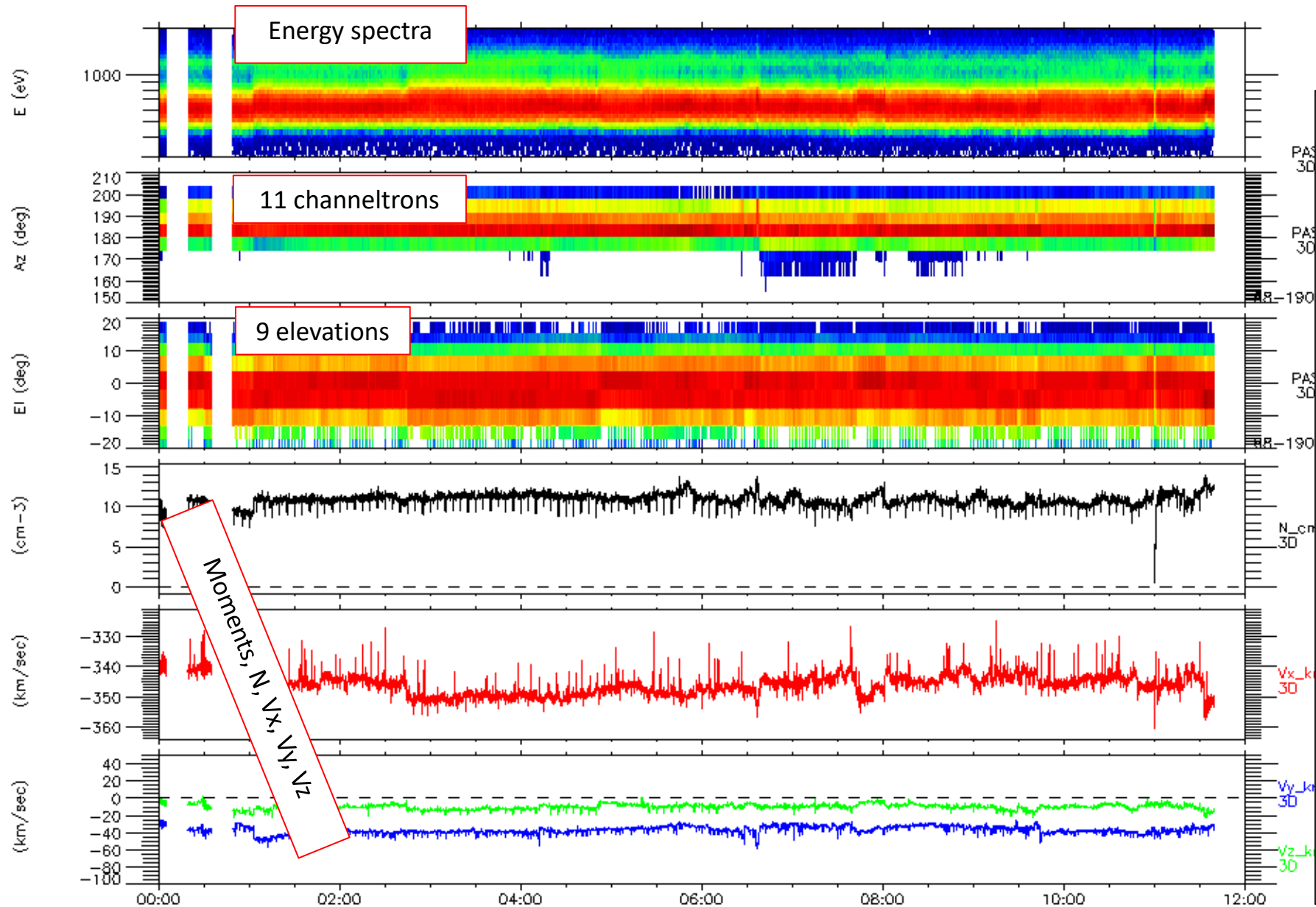
The nominal mode consists in:

- 1) Reduced distribution each 4 s
(5 ele. , 11 cha., 48 E)**
- 2) Full 3 D each 100 s for peak tracking
(9 ele. , 11 cha., 96 E)**
- 3) 8 s snapshots each 300 s. Here, we run a fast mode, at 4 Hz cadence.
(3 ele. , 11 cha., 48 E)**

Work very well. Counts are those expected.

Ground moments seem very fine.

Some details to tune



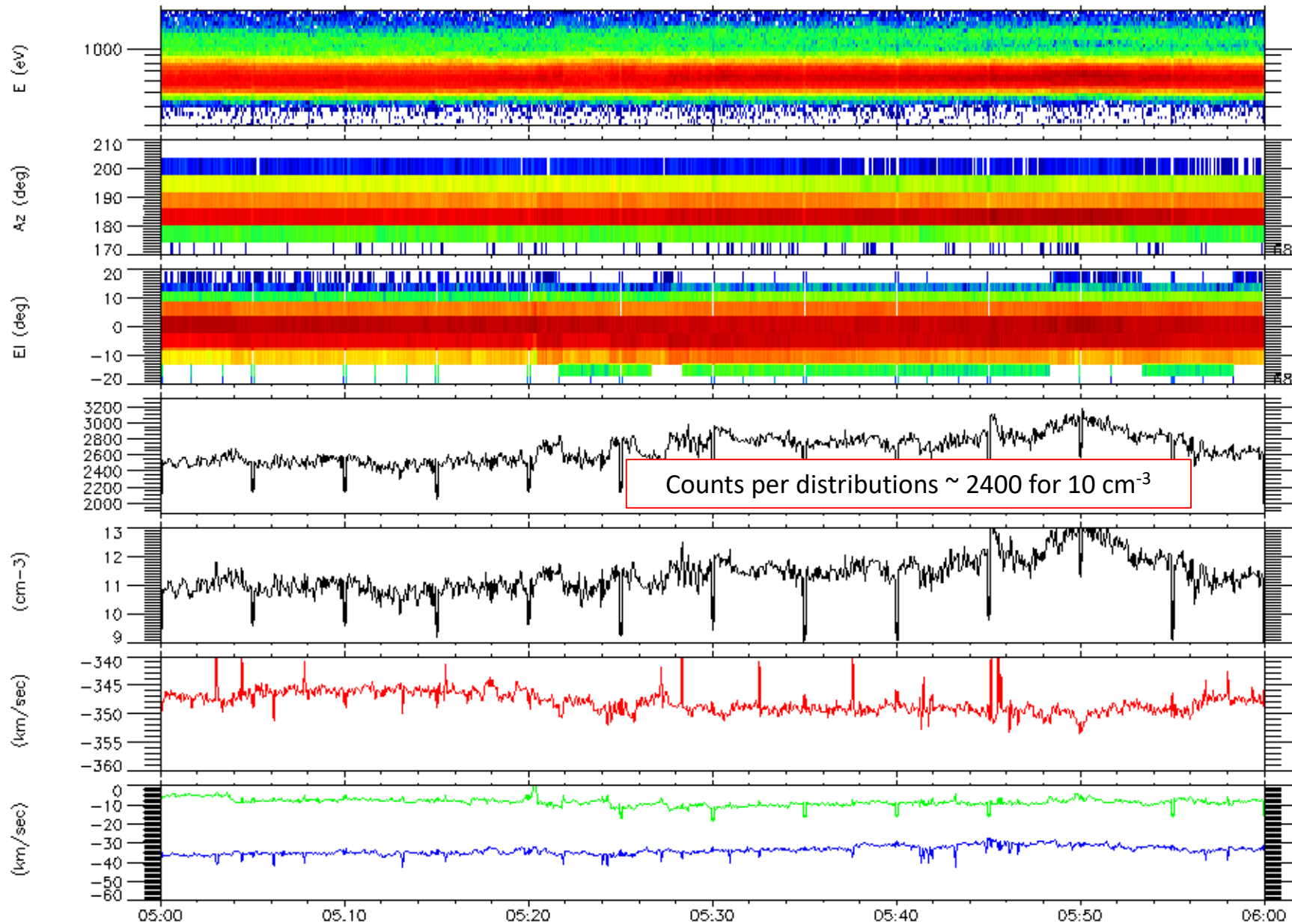
12 hours of continuous measurements.

Very close to routine operation.

- Regular spikes in density are linked to the snapshots. We use (3 ele. , 11 cha. , 48 E) and loose part of elevation. (5 ele. , 11 cha. , 32 E) will be more adapted.

- Isolated spikes in velocity are linked to isolated individual counts in phase space. We start to implement a 'cleaning' procedure.

What is the statistical accuracy of these measurements ?



Short note on statistical accuracy.

Density:

2400 counts corresponds to $\sim 10 \text{ cm}^{-3}$

$\text{Sqrt}(N)/N \sim 0.02$ (2%)

$2 \sigma \sim 100$ counts or 0.4 cm^{-3}

Typically, for 2σ :

$5 \pm 0.3 \text{ cm}^{-3}$, $10 \pm 0.4 \text{ cm}^{-3}$

$20 \pm 0.56 \text{ cm}^{-3}$, $40 \pm 0.8 \text{ cm}^{-3}$

Velocity:

$\delta V/V \sim \text{Sqrt}(T/\langle E \rangle)/\text{Sqrt}(N) \sim 0.25 \%$

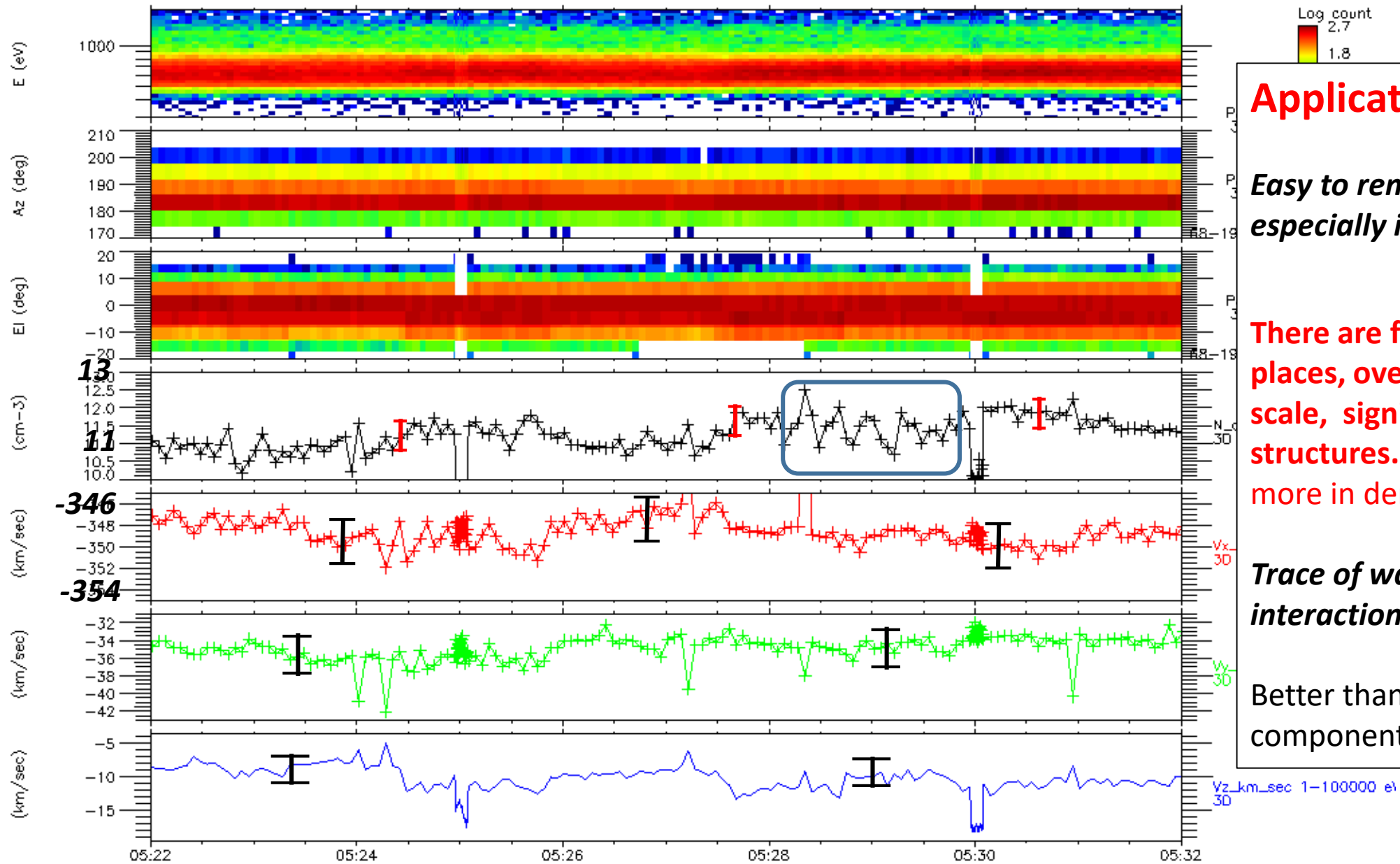
(Moore, 1998, Gersham, 2014)

Typically, for 2σ : $350 \pm 2 \text{ km/s}$

Angular accuracy: $\sim 2^\circ$

To be refined for each component.

To be done for T (but physical meaning ?)



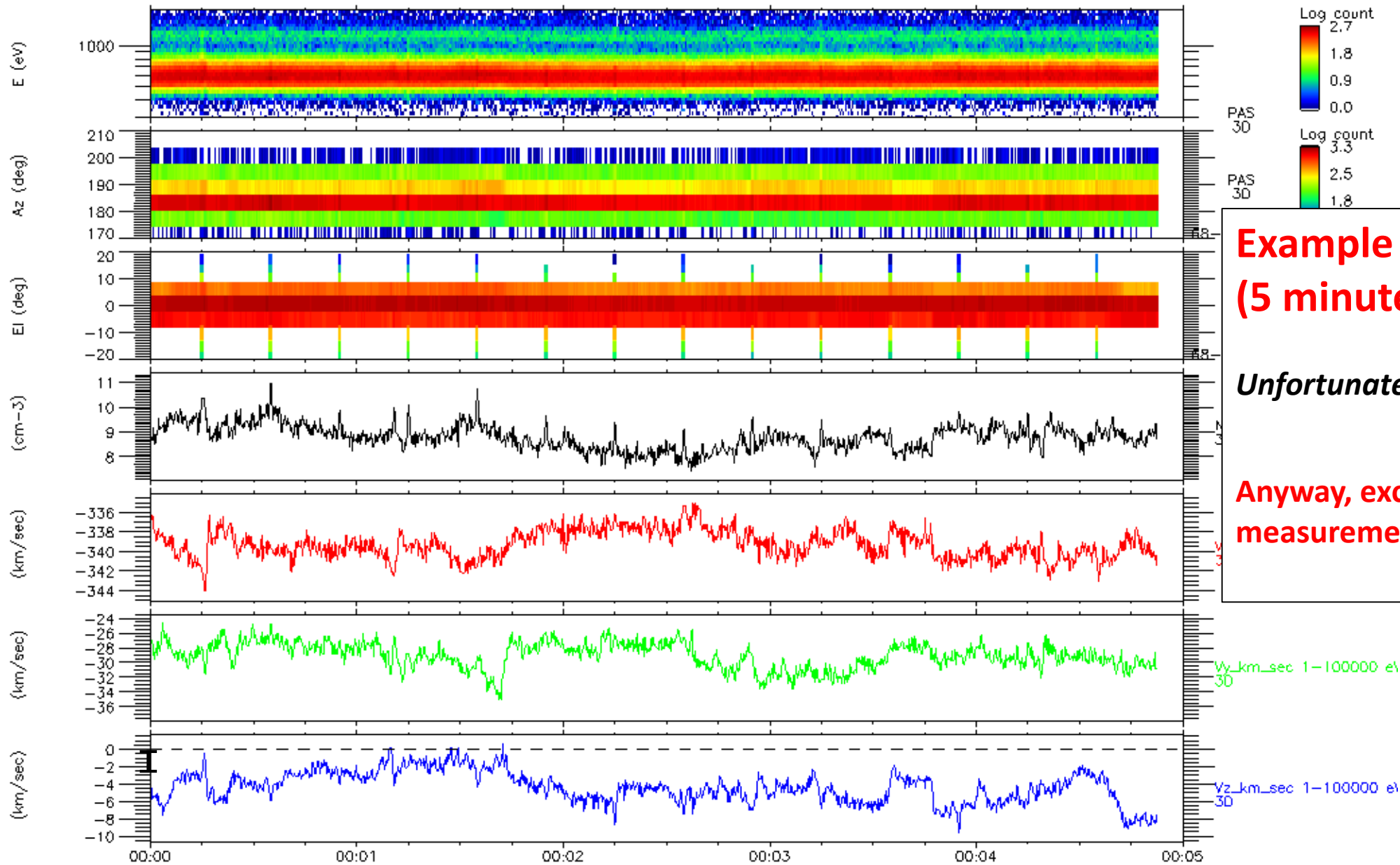
Applications.

Easy to remove single peaks, especially in V_x, V_y, V_z

There are from places to places, over minute time scale, significant wavy structures. Here, they are more in density.

Trace of wave/particle interactions ?

Better than +/- 2 km/s for V-components ?



**Example of Burst.
(5 minutes at 4 Hz)**

Unfortunately, a boring SW...

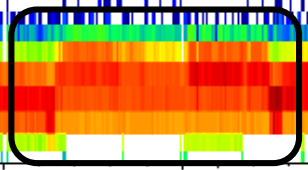
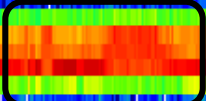
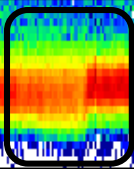
**Anyway, excellent
measurements....**

PAS

SOLAR-ORBITER

08/May/2020

E-disp.

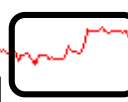


Ele.jump

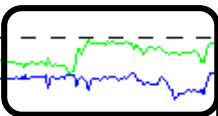
N-Fluc.



V-Fluc.



Flow-devia.

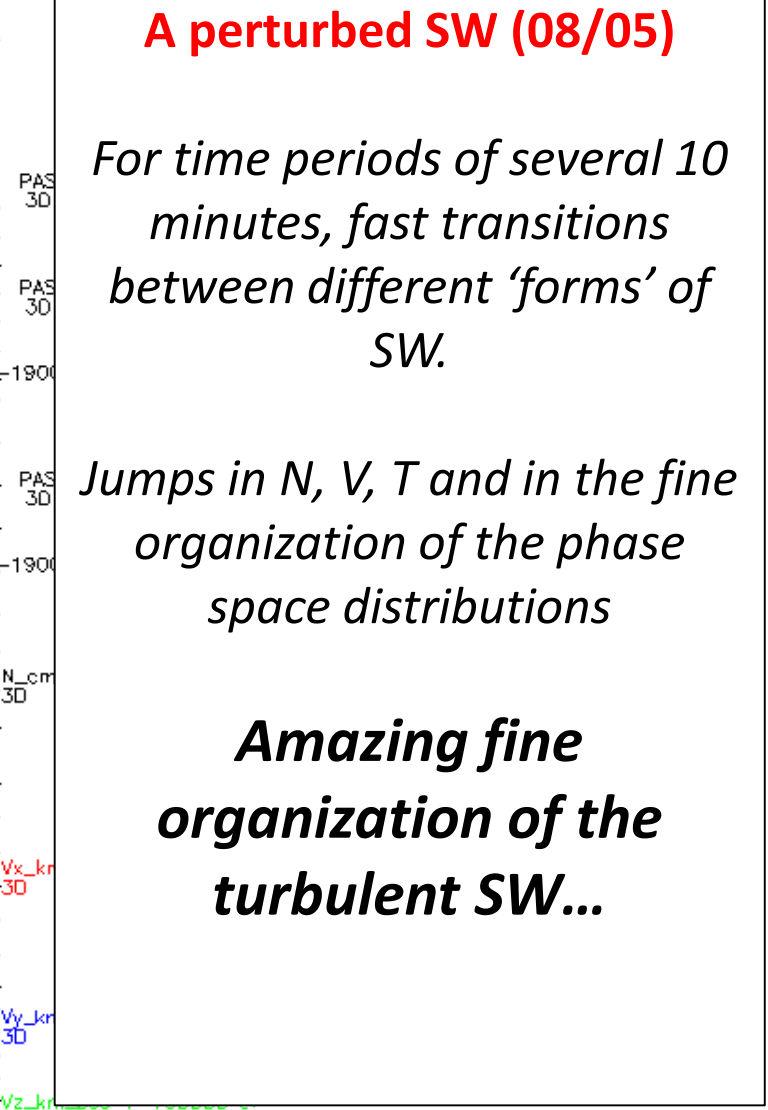


A perturbed SW (08/05)

For time periods of several 10 minutes, fast transitions between different 'forms' of SW.

Jumps in N, V, T and in the fine organization of the phase space distributions

Amazing fine organization of the turbulent SW...

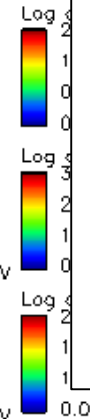
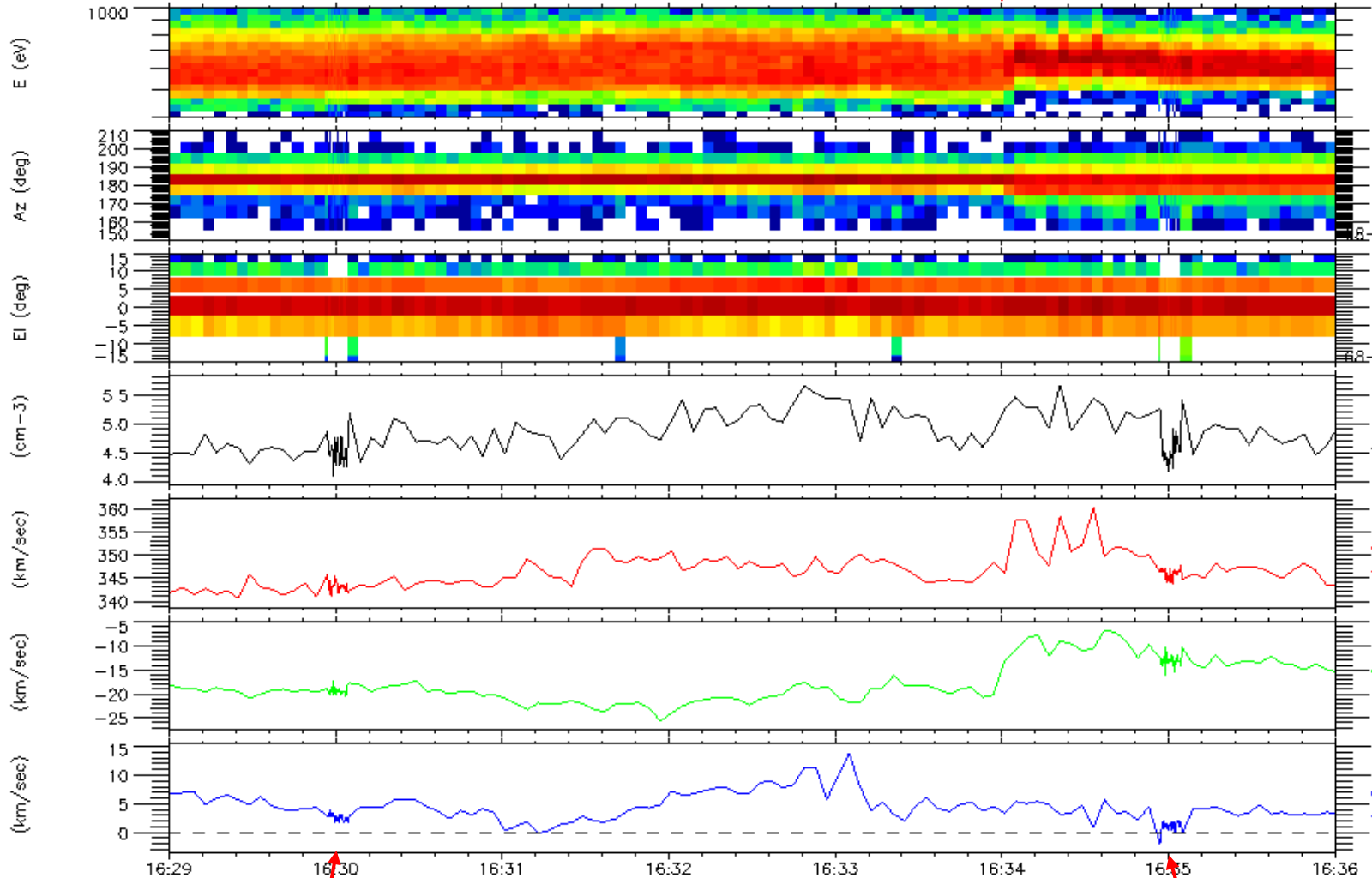


PAS SOLAR-ORBITER Transition 15/Apr/2020

SW state 1

SW state 2

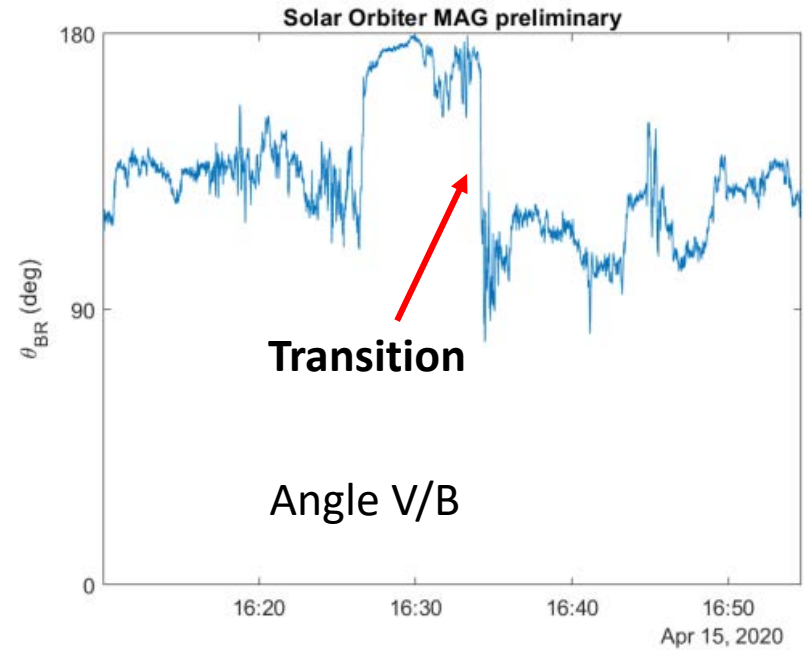
Transition



Observation of a SW perturbation. 15/04

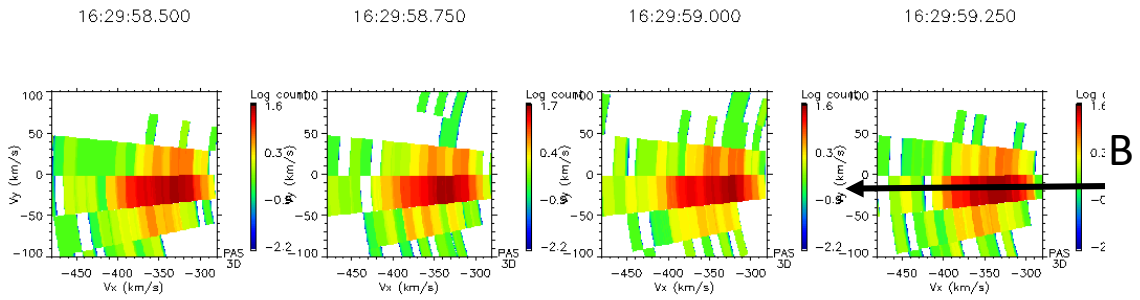
Sudden change of the SW parameters (anisotropy).

Correspond also to a perturbation seen with MAG

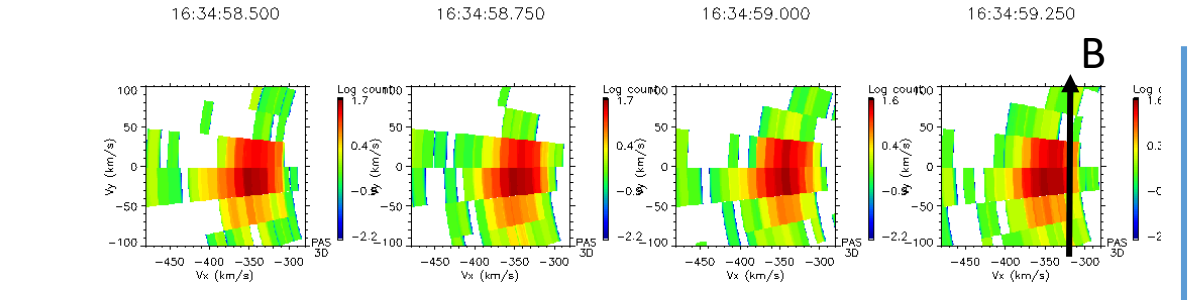
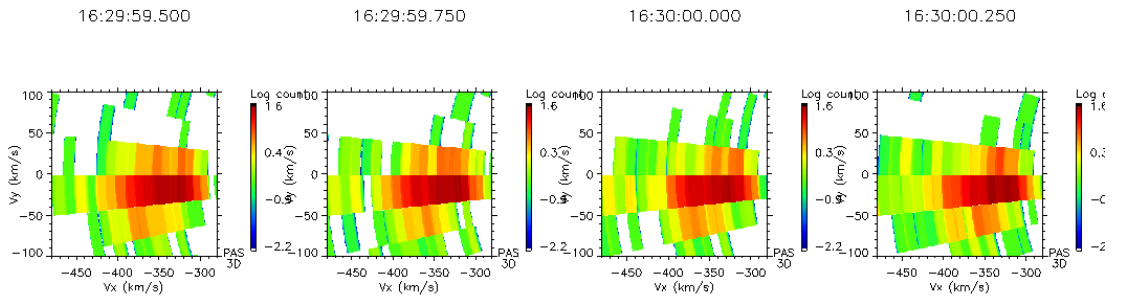


Snapshot 1 (4 Hz resolution)

Snapshot 2



Snapshot 1. Center 16:30 . 4 Hz



Snapshot 2. (5 minutes later) Center 16:35 . Again 4 Hz



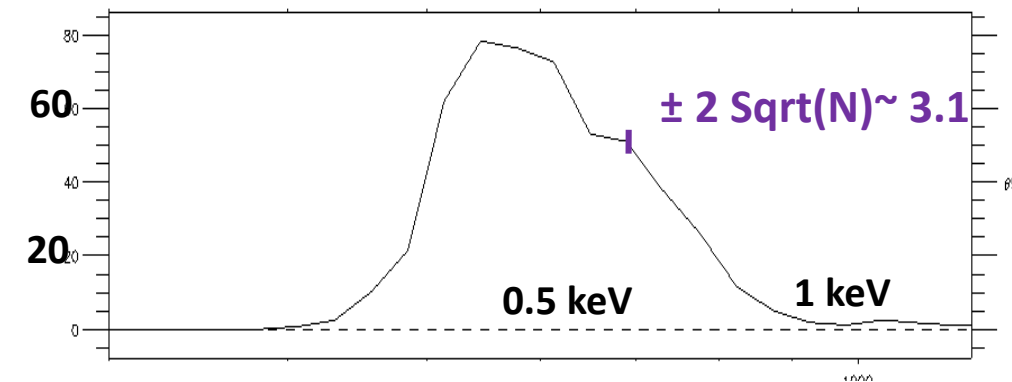
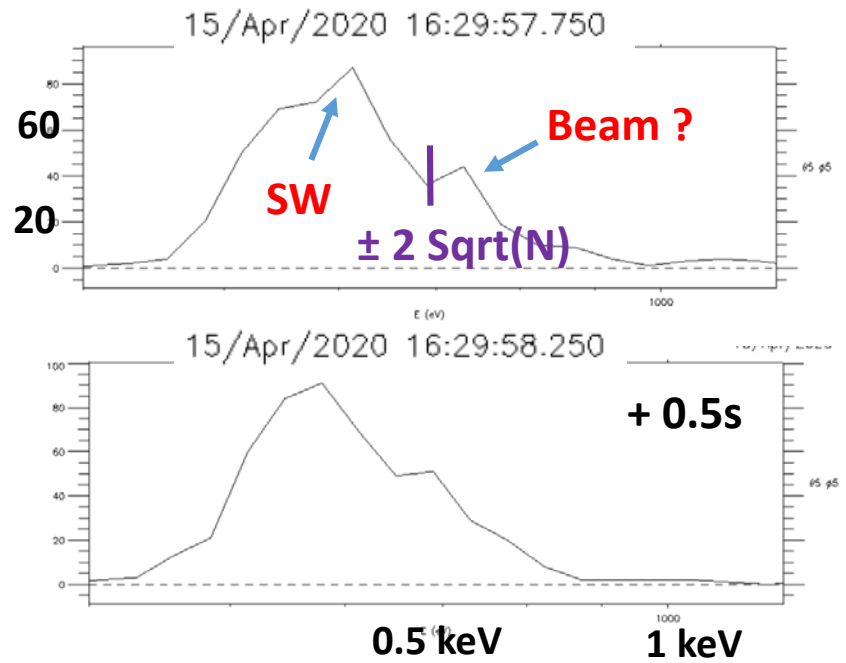
To increase statistics: average over 20 distributions (almost the whole snapshot).

Produced by CLWeb

Same n, same V but a different solar wind.

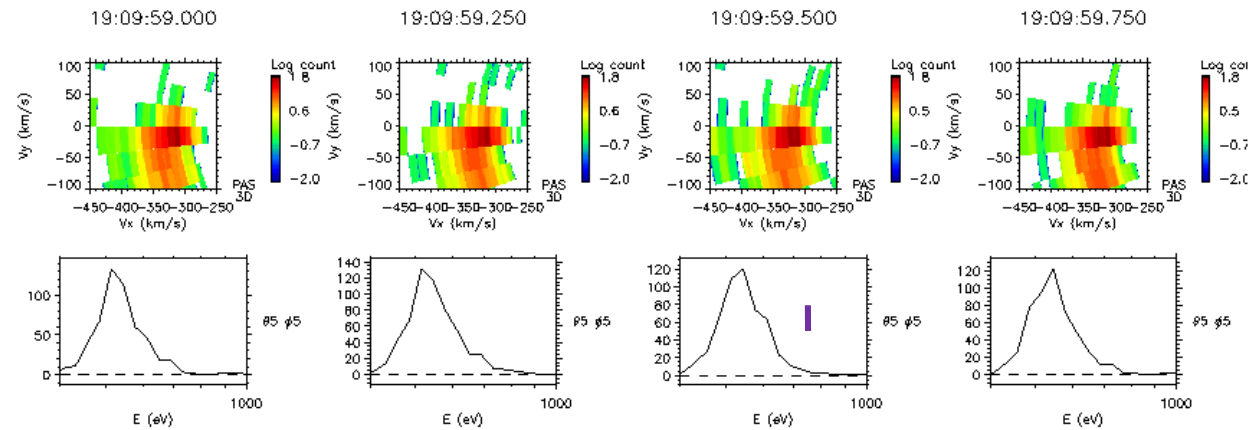
Possible fast evolution of beam... (1s time scale)

Is there really a beam ?



More than 1000 counts to define the beam. When normalized, $2\sigma \sim 3.1$

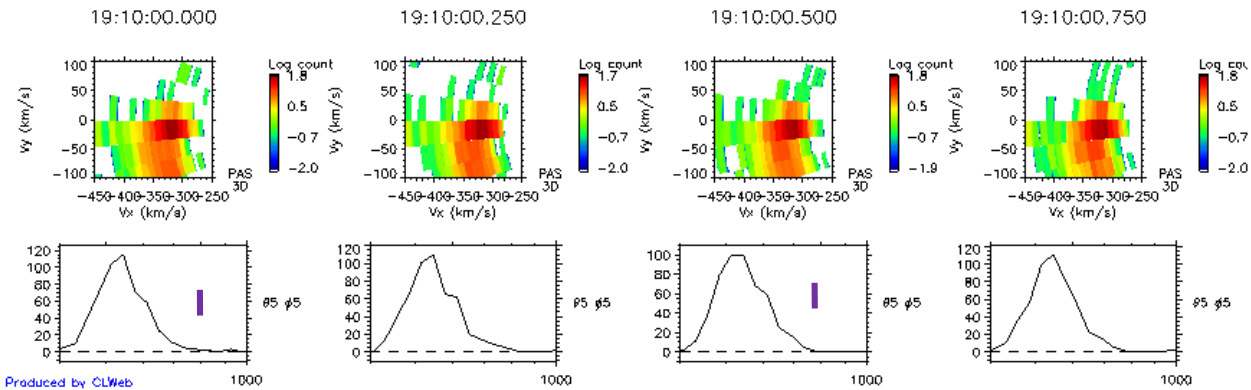
The beam is real but difficult to follow its dynamics...



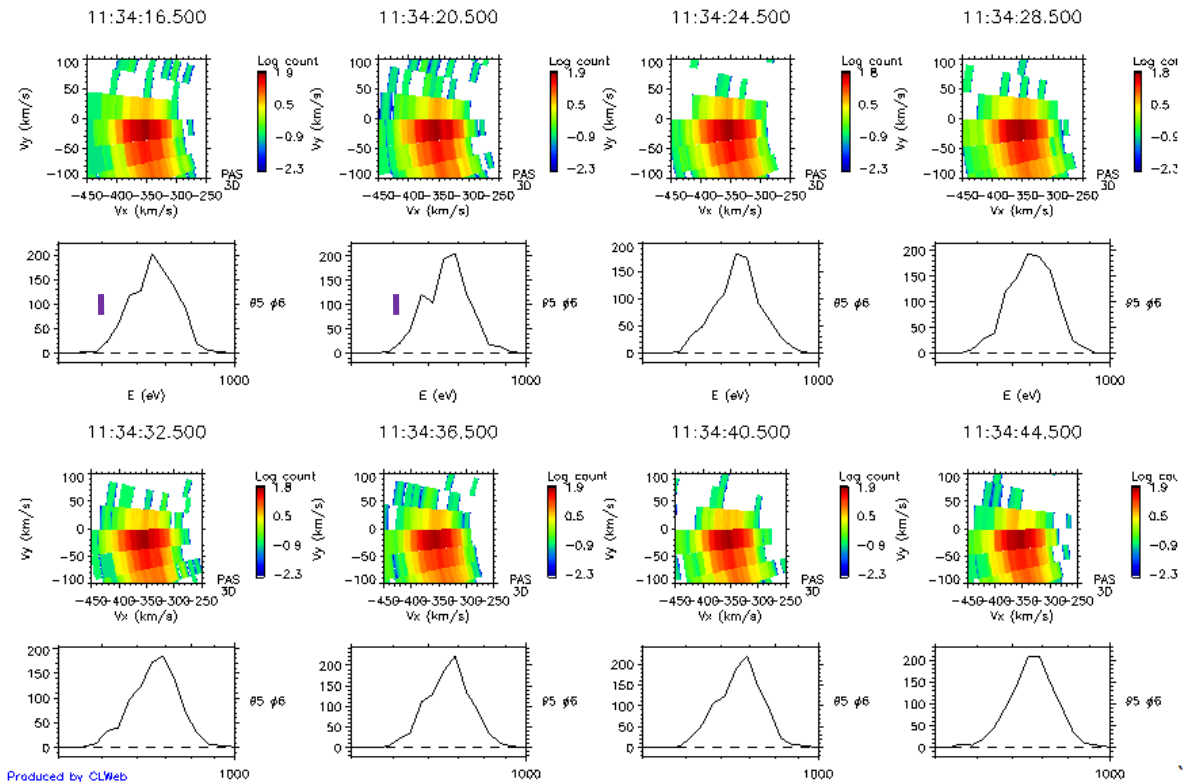
Example of marginally significant fluctuations.

Secondary beam dynamics

(Thanks to Tristan Hoellinger)



Secondary beam at 500 eV ?



Secondary beam at 600 eV ?

**Marginal at 2 sigma for single spectrum
Beam existence should be confirmed by summation.**

Try 1 elevations ?

Better statistical tools (gaussian average)

Things will better closer to Sun.

New observations (?)

'Double' SW

(08/05)

Observation of 2 peaks of equivalent density.

First separated in azimuth (12-15°) then in energy (200 eV).

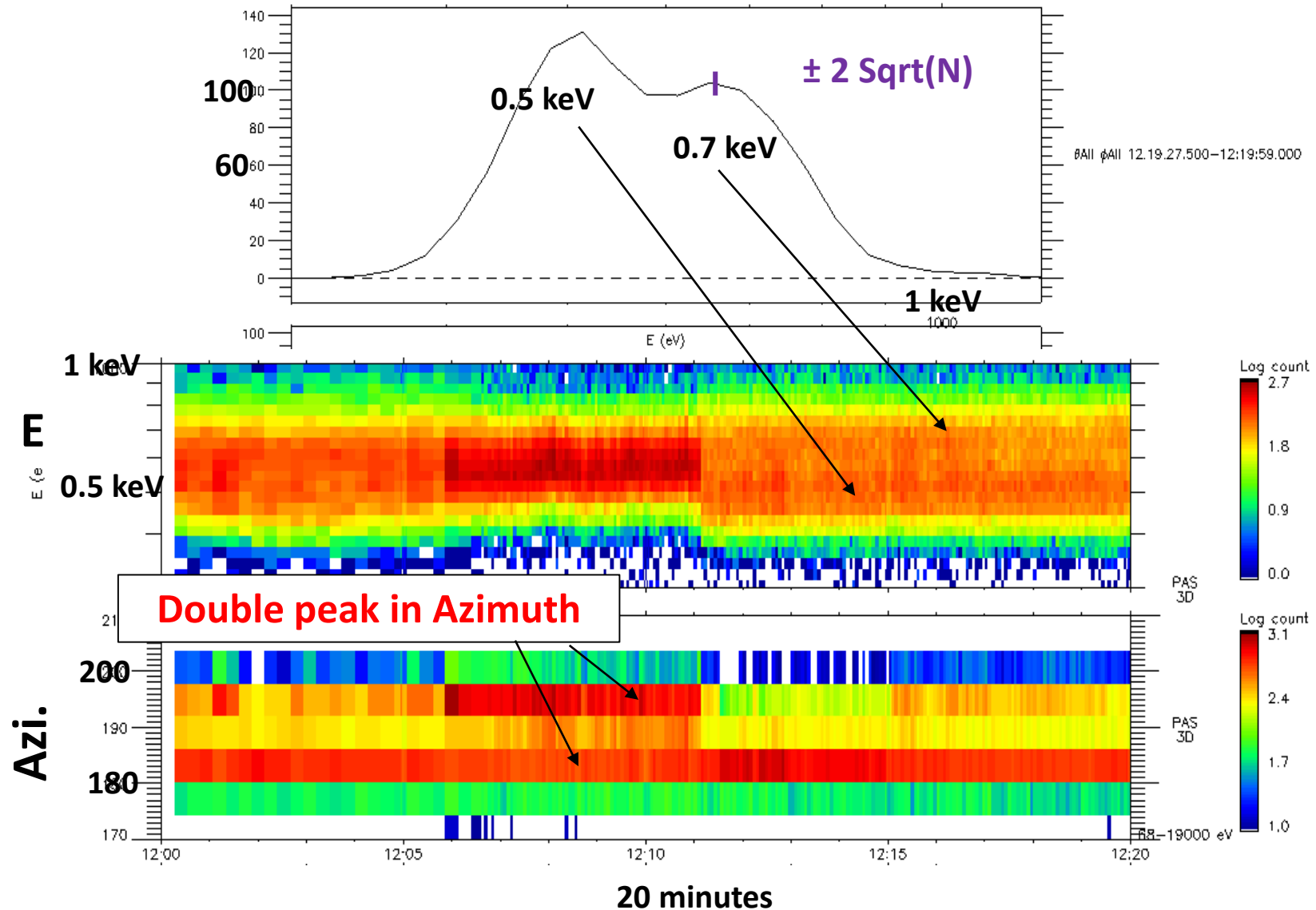
Very sharp transition (~4 s)

What does it mean at kinetic scales ?

Average over 15 distributions

Double peak in energy

500



CONCLUSION

PAS works very well !

Minor defaults that are easy to correct.

'Ghosts' due to azimuthal asymmetry (come from an obligation of the design...).

Scientific performances are at 'rendezvous'.

When density $> 8-10 \text{ cm}^{-3}$ (typically, 150 counts on central peak) and some adaptations (sum on a few distributions) , the statistics is sufficient for the study of kinetic processes (beams dynamics) at subsecond resolution.

With PAS, a new chapter on SW dynamics, at kinetic scales...

Thanks to all the team and especially to Andrei